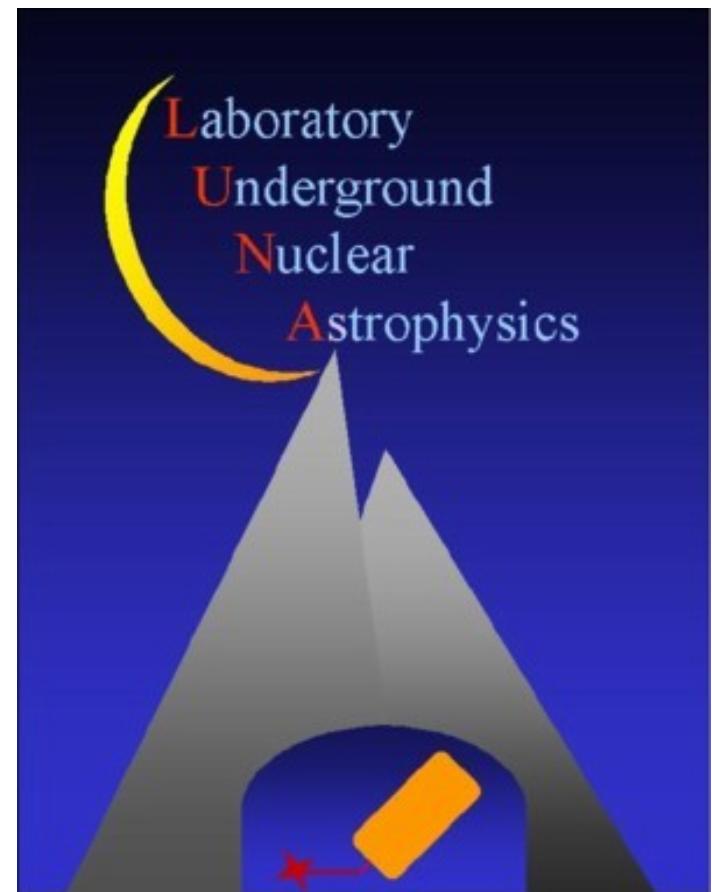


# Study of the BBN reaction D( $\alpha,\gamma$ ) $^6\text{Li}$ deep underground in LUNA

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(INFN, Sez. di Padova)  
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19.07.2010

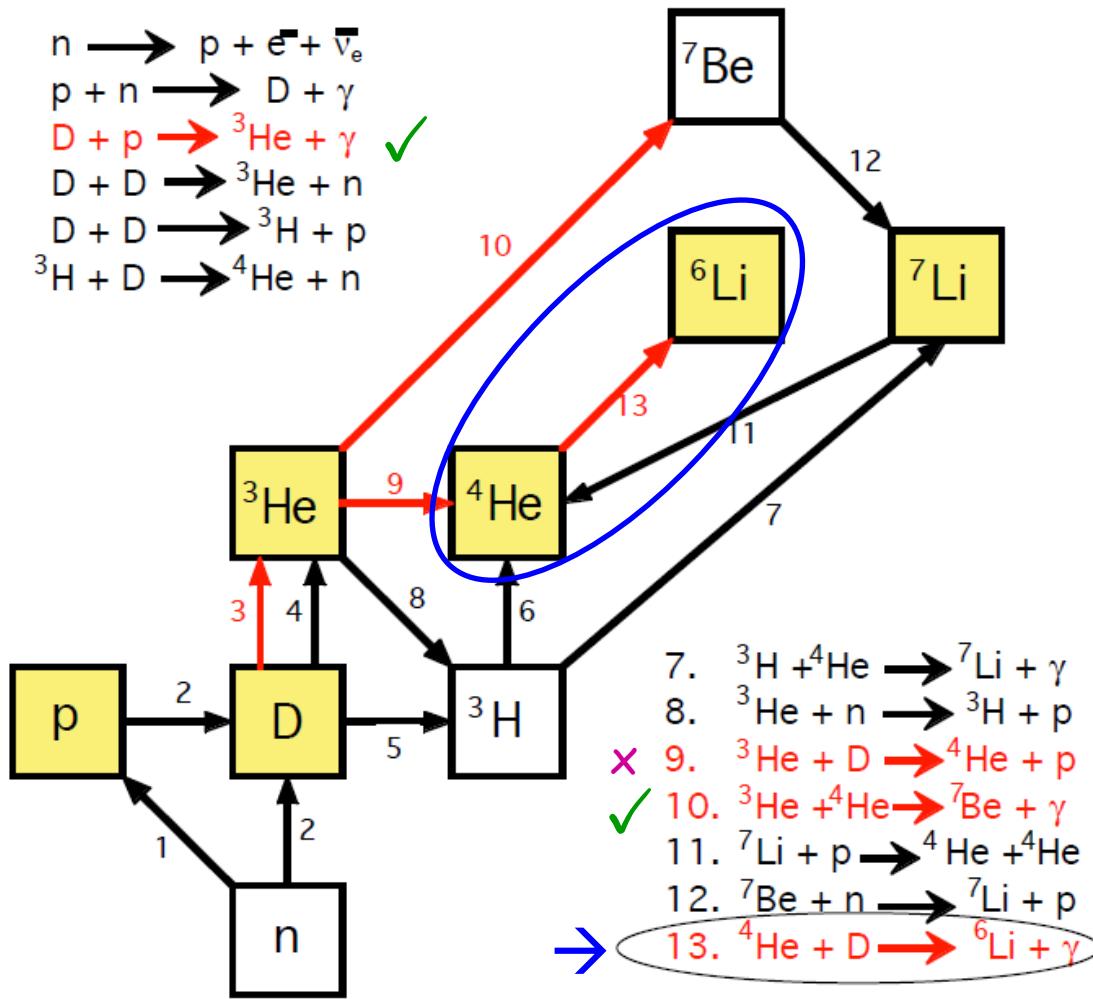
Symposium on  
***Nuclei in the Cosmos XI***  
*Heidelberg*  
19 - 23 July 2010



# The BBN reaction network

Schematic BBN network:

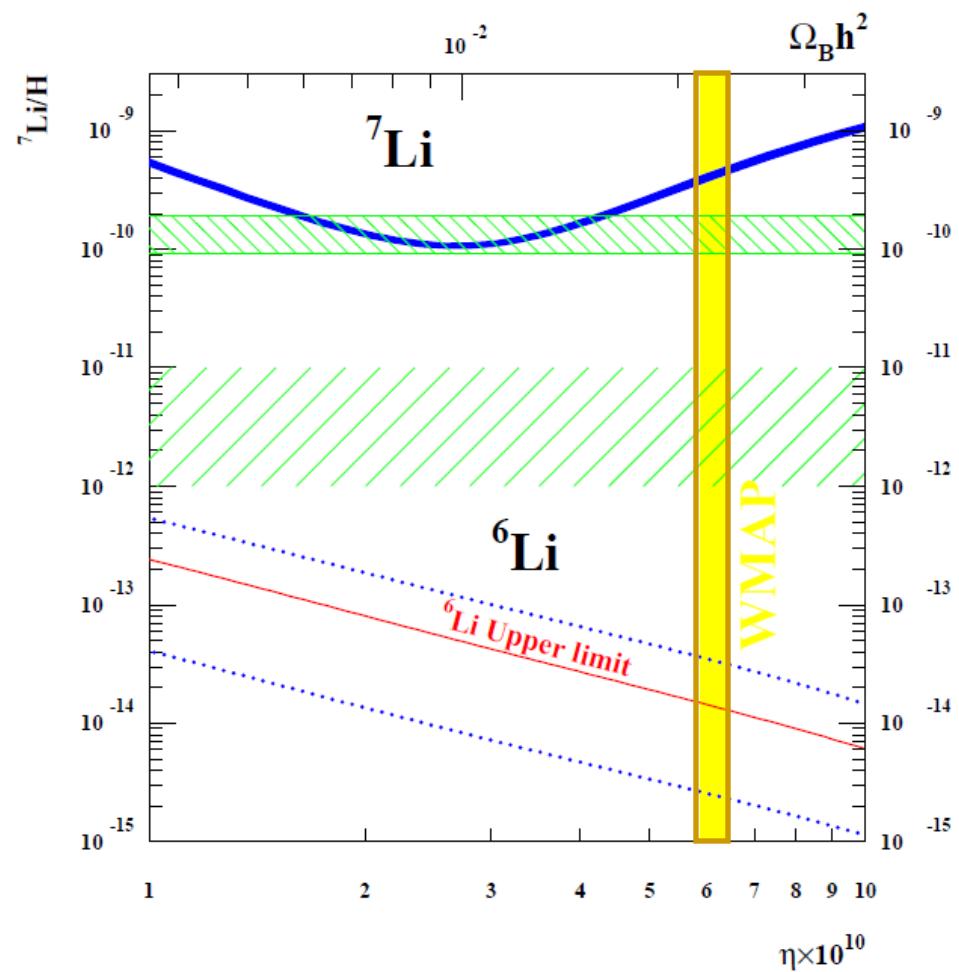
1.  $n \rightarrow p + e^- + \bar{\nu}_e$
2.  $p + n \rightarrow D + \gamma$
3.  $D + p \rightarrow {}^3\text{He} + \gamma$  ✓
4.  $D + D \rightarrow {}^3\text{He} + n$
5.  $D + D \rightarrow {}^3\text{H} + p$
6.  ${}^3\text{H} + D \rightarrow {}^4\text{He} + n$



- apart for Helium, all the other nuclides are sensitive to the nuclear reaction network
- good agreement between calculations and observations for  $D$ ,  ${}^{3,4}\text{He}$
- problems for  ${}^{6,7}\text{Li}$ !
- Already measured by LUNA:
  - $D(p,\gamma){}^3\text{He}$
  - ${}^3\text{He}(\alpha,\gamma){}^7\text{Be}$
  - ( $D$ ,  ${}^7\text{Li}$  abundance)
- Currently investigated:
  - **$D(\alpha,\gamma){}^6\text{Li}$**
  - ( ${}^6\text{Li}$  abundance)
- Spin off:
  - $D({}^3\text{He},p){}^4\text{He}$
  - ( ${}^3\text{He}$  abundance)

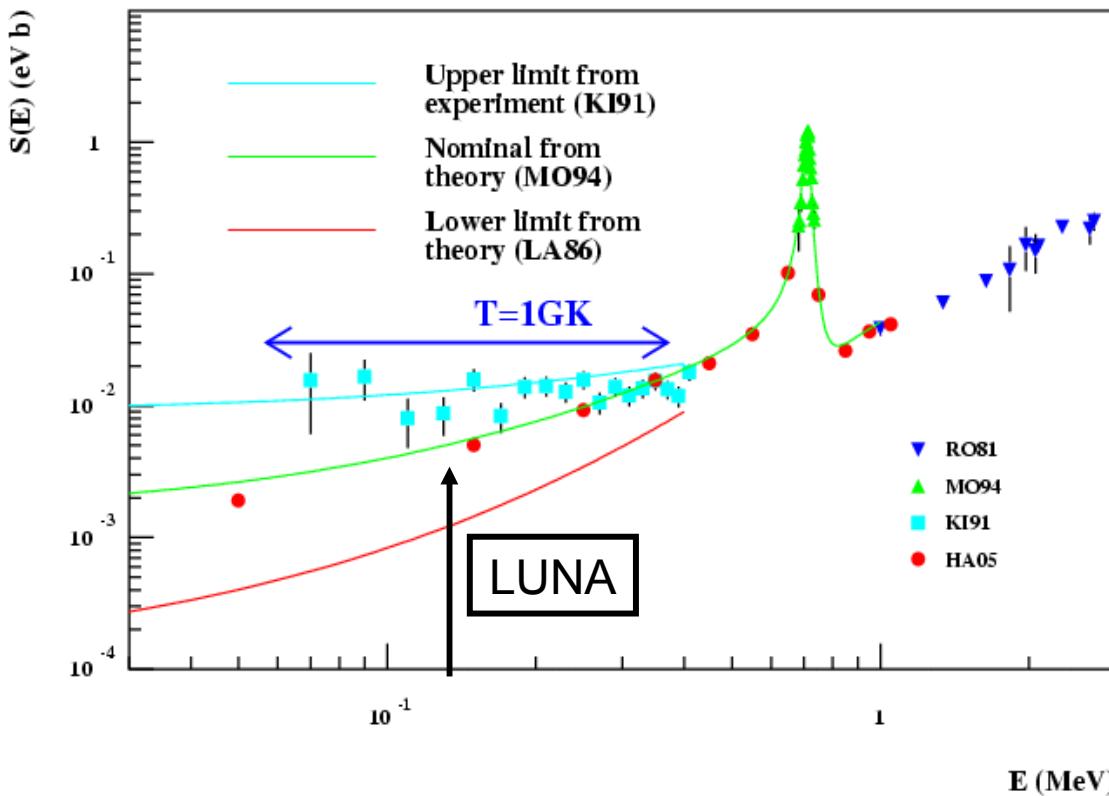
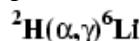
# BBN: ${}^6/{}^7\text{Li}$ abundance problem

- ${}^6\text{Li}$  detected in metal poor stars (green bars) is unexpectedly large compared to the NACRE BBN predictions (blue dashed lines)
- upper limit from Coulomb Dissociation experiment of  ${}^6\text{Li}$  @ GSI (indirect measurement) (red line)
- baryon-to-photon ratio  $\eta$  fixed by WMAP

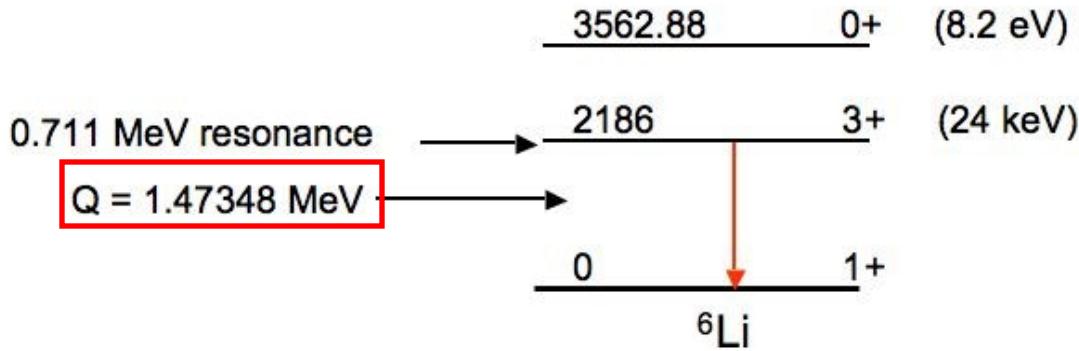


# Cross section of the D+ $\alpha$ reaction

- theoretical estimates  $\rightarrow$  differ by >1 order of magnitude
- no direct measurements < 650 keV (resonance: 711 keV)
- cross section very low ( $\sigma(E_{\text{c.m.}}) = S(E_{\text{c.m.}}) / E_{\text{c.m.}} \cdot e^{-2\pi\eta}$ )
- big uncertainty extrapolating S-factor ( $\rightarrow S_0$ )

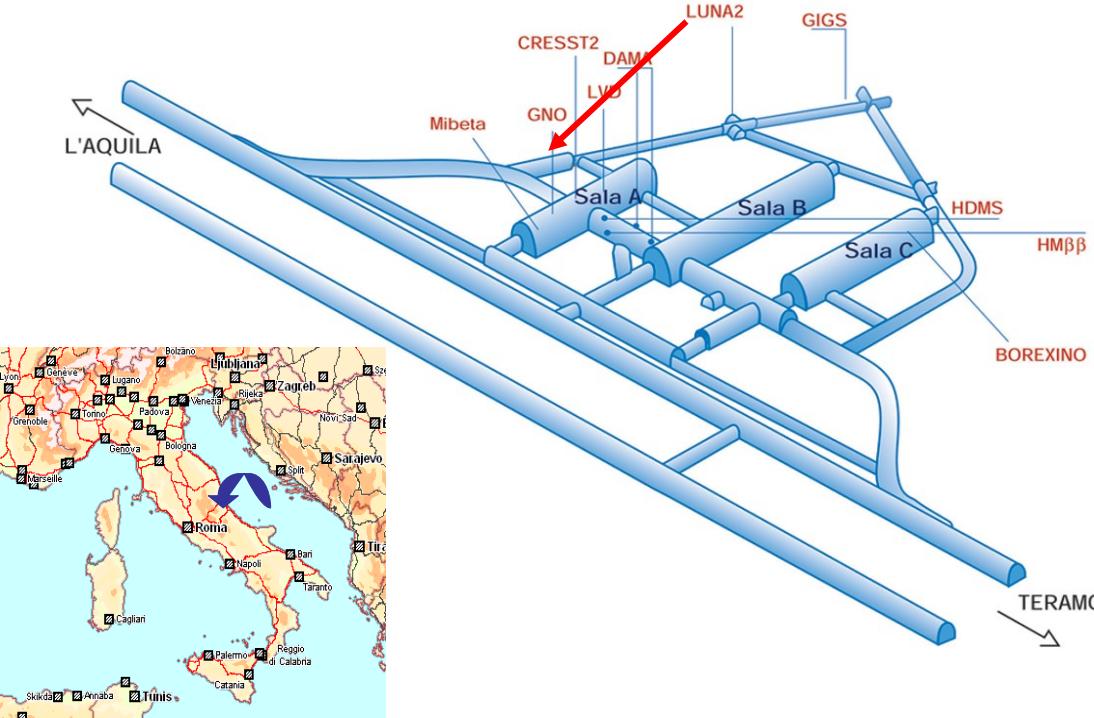


- direct measurements desirable in the **BBN energy window** to test models  
 $E_{\text{BBN}} = 50 - 300$  keV
- cross section:  
 $\sigma(E_{\text{BBN}})$ : 10 fb – pb
- LUNA, D+ $\alpha$ :  
 $E_\alpha \leq 400$  keV  $\rightarrow$   
 $E_{\text{c.m.}} \approx 130$  keV



- $E_\gamma = Q + E_{\text{cm}} + \Delta E_{\text{Dopp}} - \Delta E_{\text{rec}}$   
 $= \text{RoI} = 1580-1630 \text{ keV}$ ,  
 in-beam  $\gamma$ -spectroscopy
  - direct measurement possible  
 @ LUNA
  - LNGS: reduction of  
 $\mu$  by  $10^6$ ,  $n$  by  $10^3$
- talk **H. Costantini**  
 (Session 4: Stars,  
 Mon., 19<sup>th</sup>, 17:00 h)

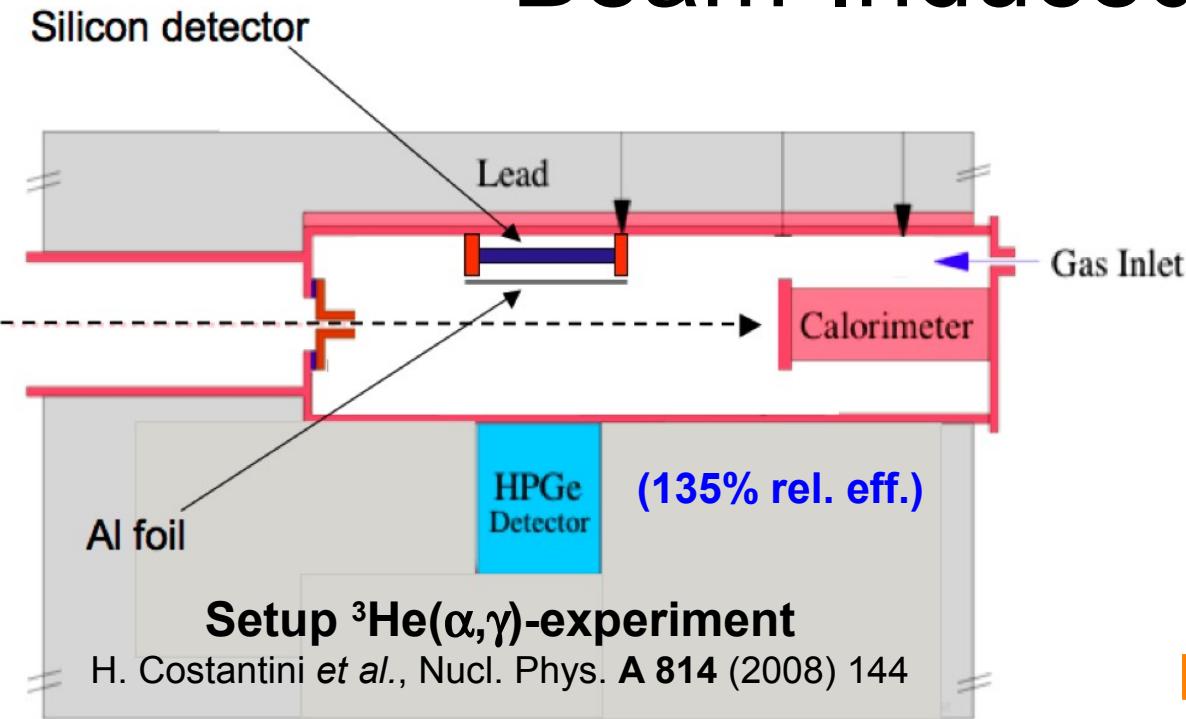
## Laboratory for Underground Nuclear Astrophysics



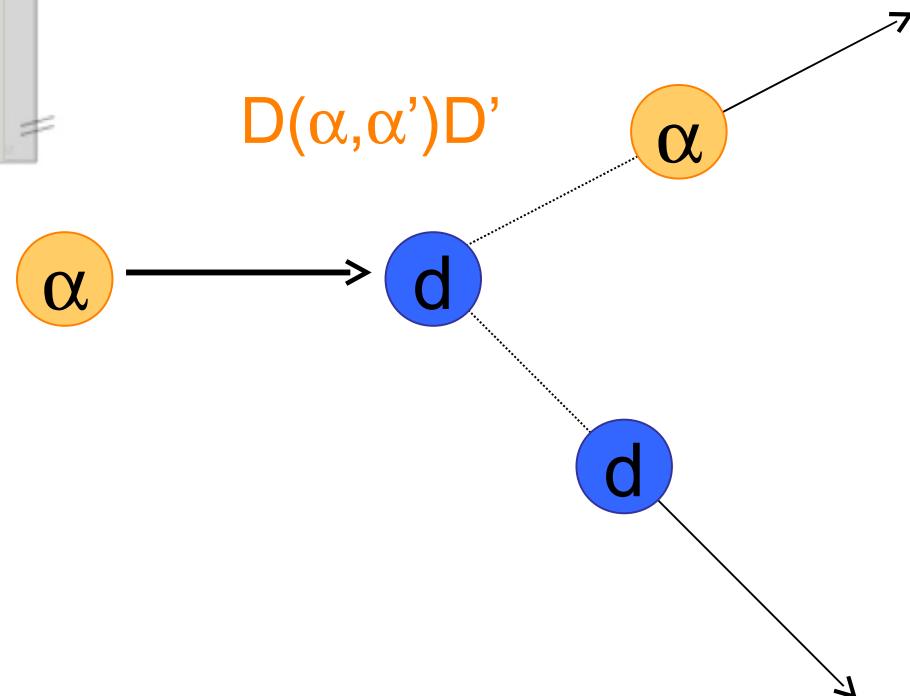
**Laboratori Nazionali  
del Gran Sasso (LNGS)**  
 (shielding:  
**1400 m dolomite rock**  
 → **3800 m w.e.**)



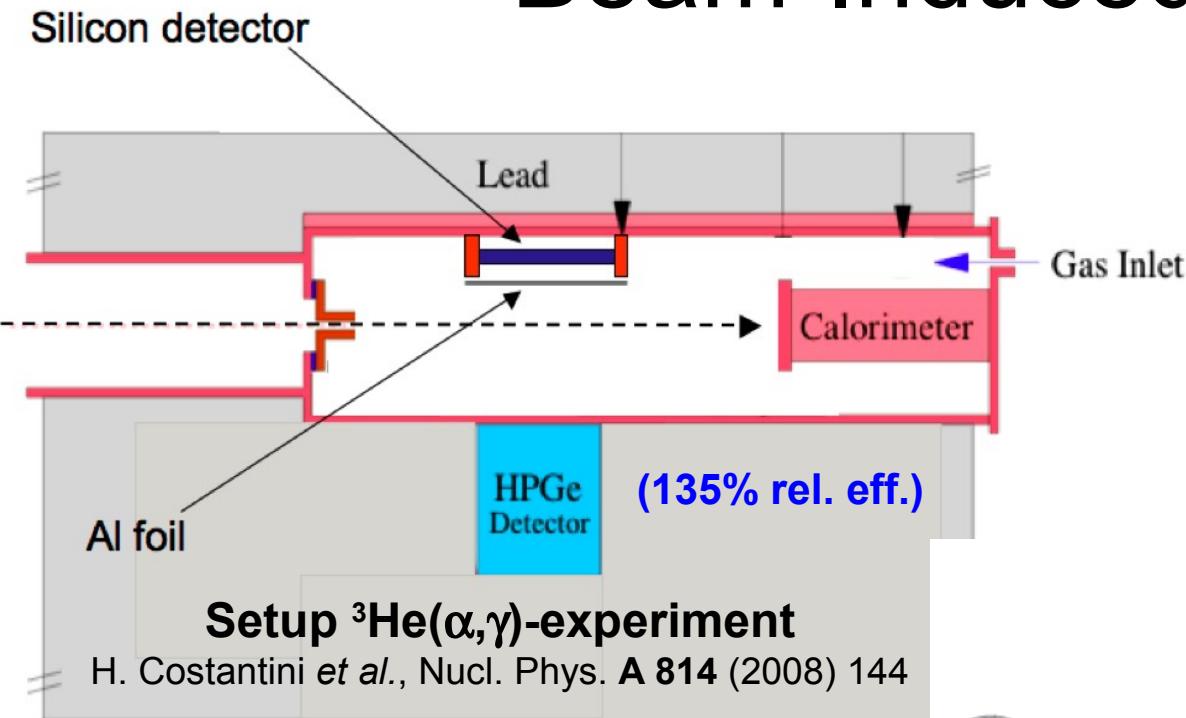
# Beam-Induced Background (BIB)



**${}^4\text{He}^+$ -beam**  
( $E_{\max} = 400 \text{ keV}$ ,  
 $I_{\max} \approx 250 \mu\text{A}$ )



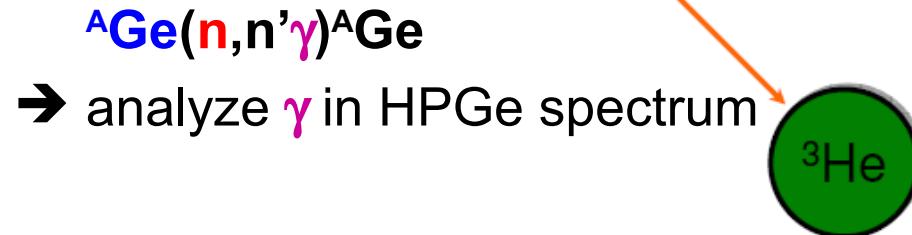
# Beam-Induced Background (BIB)



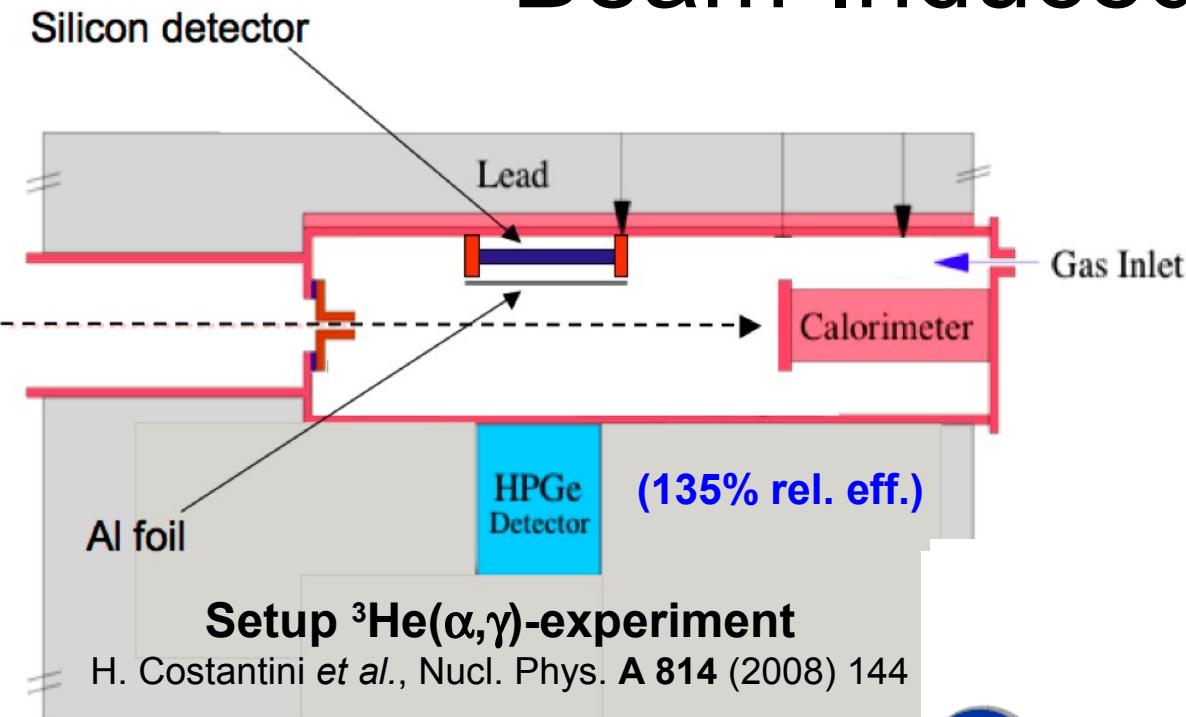
**${}^4\text{He}^+$ -beam**  
( $E_{\text{max}} = 400 \text{ keV}$ ,  
 $I_{\text{max}} \approx 250 \mu\text{A}$ )

$$\sigma_{D(d,n){}^3\text{He}}$$

$Q = 3.269 \text{ MeV}$   
 $E_n = 2.450 \text{ MeV}$   
 $E_{{}^3\text{He}} = 0.820 \text{ MeV}$



# Beam-Induced Background (BIB)

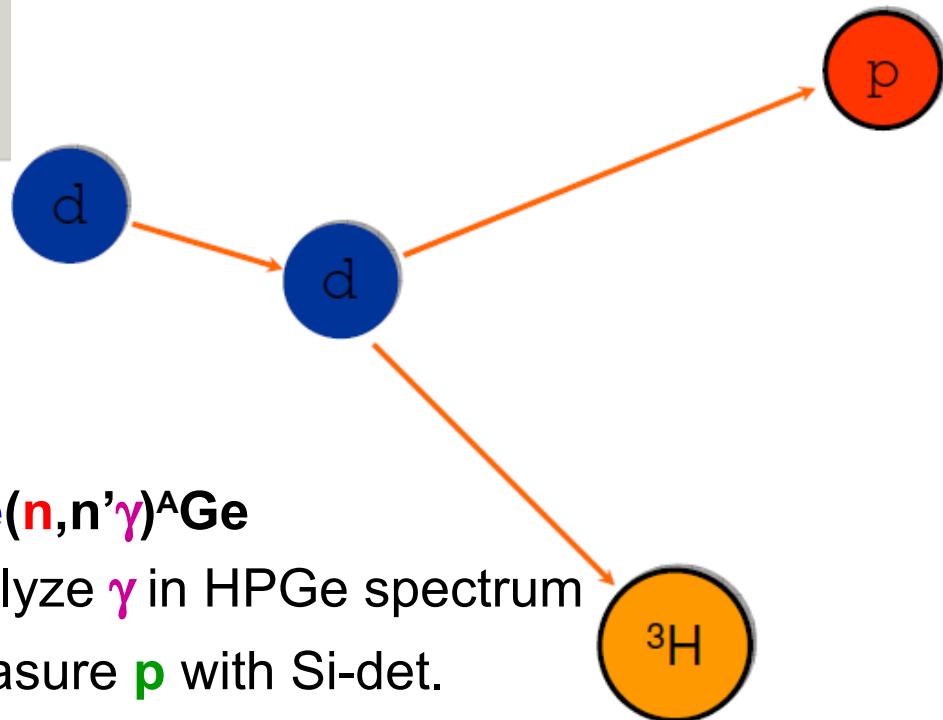


**${}^4\text{He}^+$ -beam**  
( $E_{\max} = 400 \text{ keV}$ ,  
 $I_{\max} \approx 250 \mu\text{A}$ )

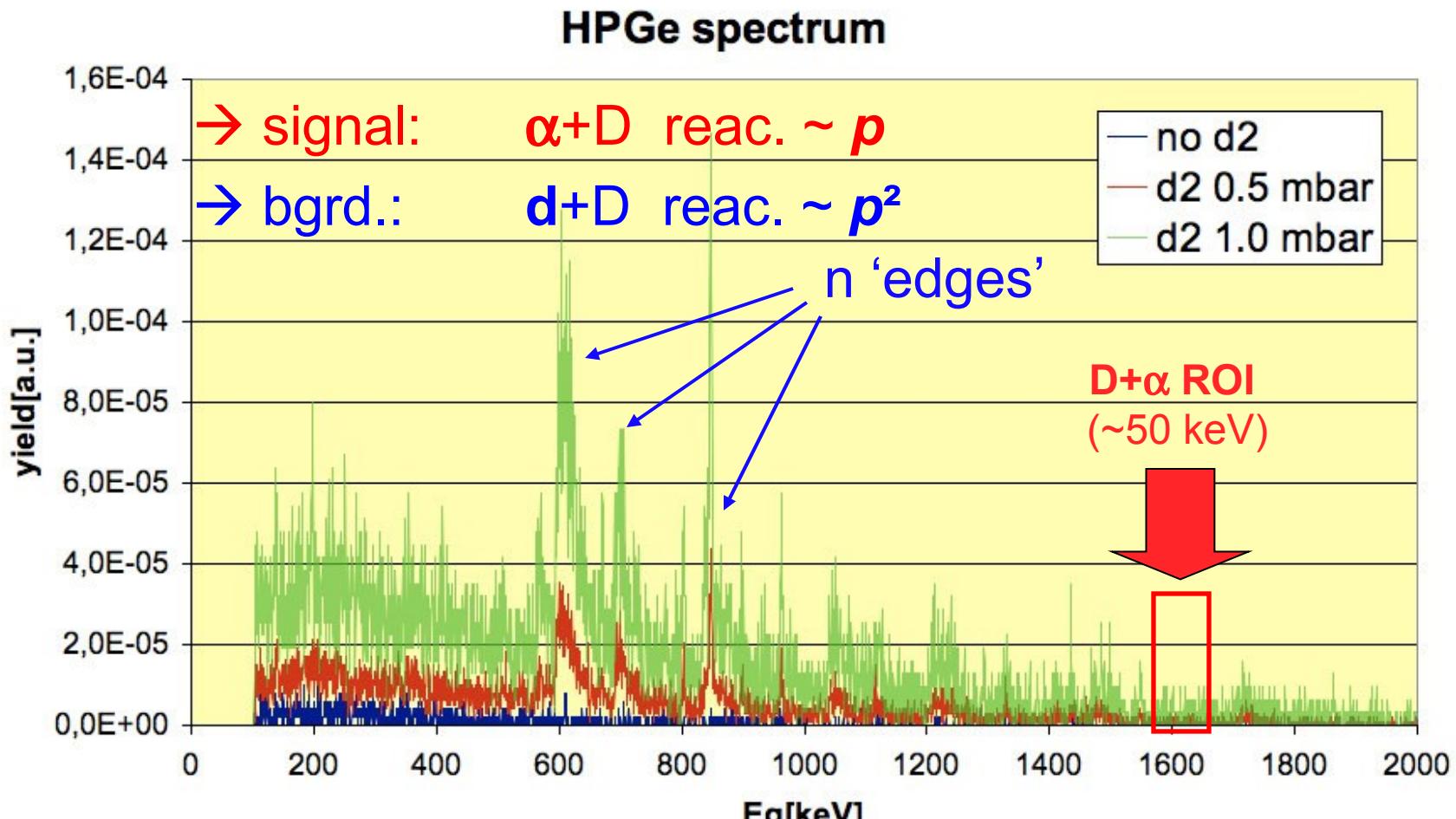
$$\sigma_{D(d,n){}^3\text{He}} \equiv \sigma_{D(d,p)t}$$

$Q = 3.269 \text{ MeV}$	$Q = 4.033 \text{ MeV}$
$E_n = 2.450 \text{ MeV}$	$E_p = 3.022 \text{ MeV}$
$E_{{}^3\text{He}} = 0.820 \text{ MeV}$	$E_t = 1.011 \text{ MeV}$

- ${}^A\text{Ge}(n,n'\gamma){}^A\text{Ge}$
- analyze  $\gamma$  in HPGe spectrum
  - measure  $p$  with Si-det.



# $^{A\text{Ge}}(\text{n}, \text{n}'\gamma) \rightarrow \text{n-fluence}$

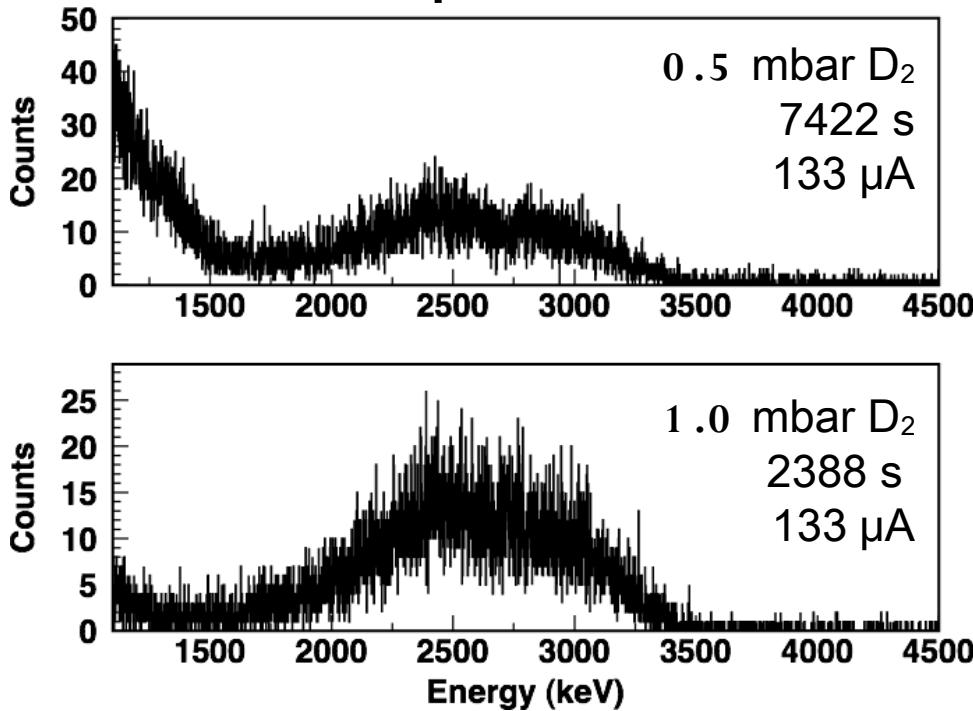


$$n / \text{cm}^2 = 300 * N_{\gamma}(693 \text{ keV}) / V_{\text{HPGe}} [\text{cm}^3]$$

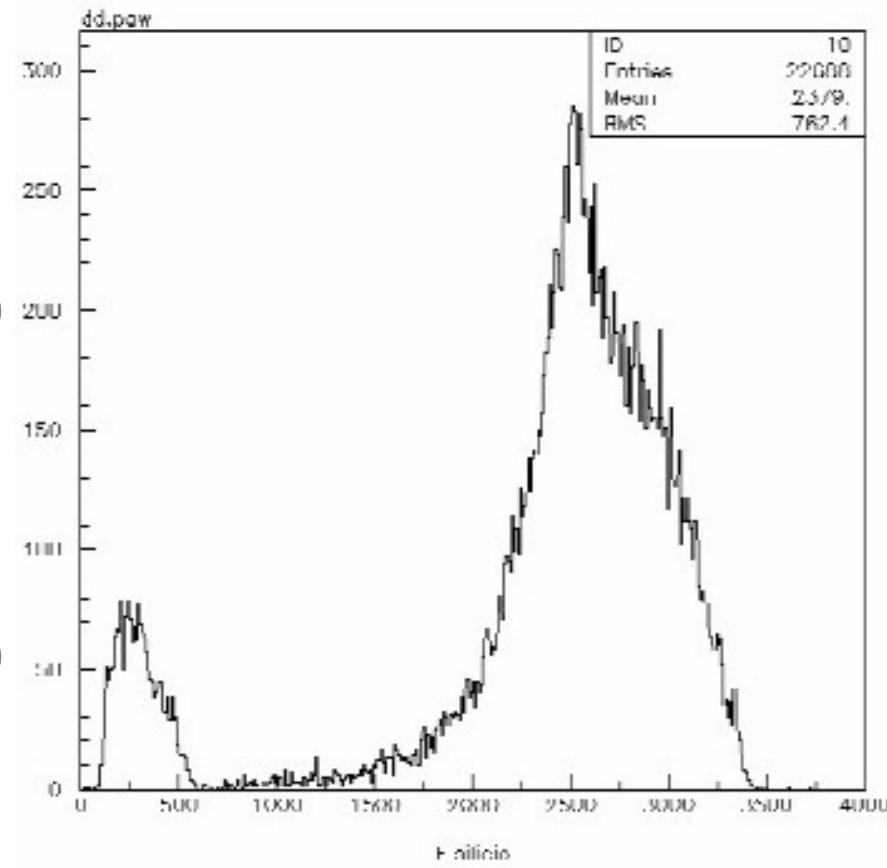
# Si-detectors: D(d,p) → n-flux

→ measure p from D+d-reaction with Si-det.:

Experiment

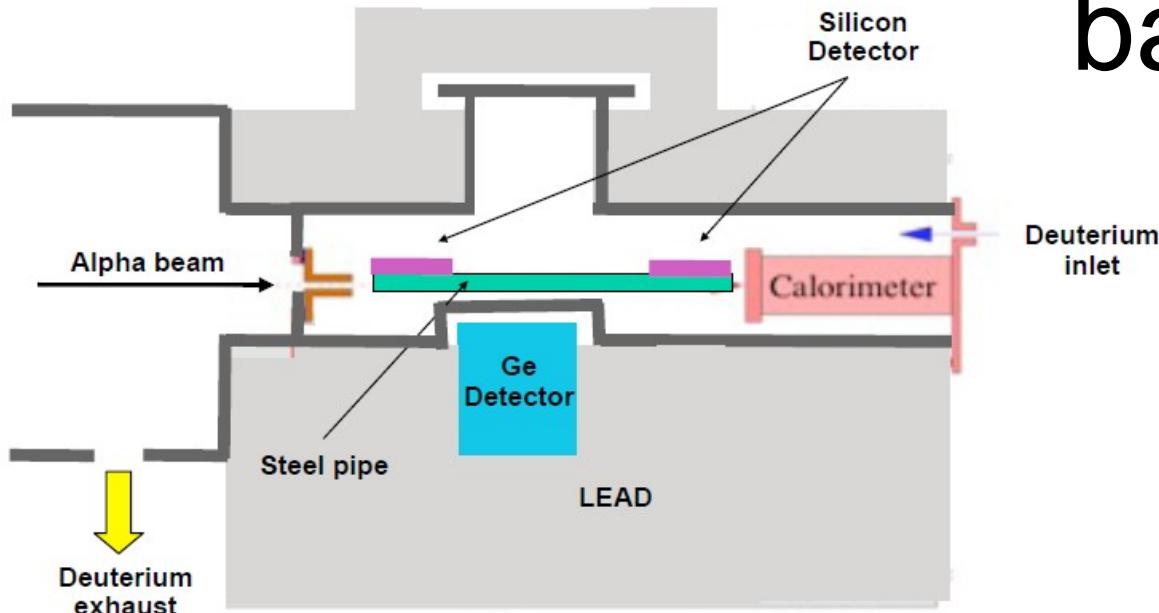


MC Simulation

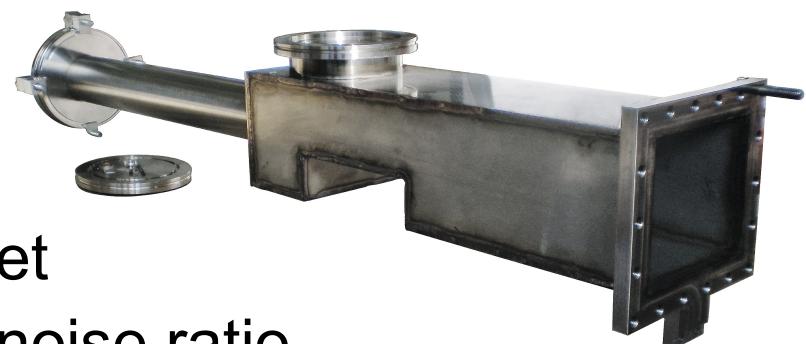


- Monte-Carlo simulations confirmed ✓
- both methods coincide ✓

# Reduction of beam-induced background

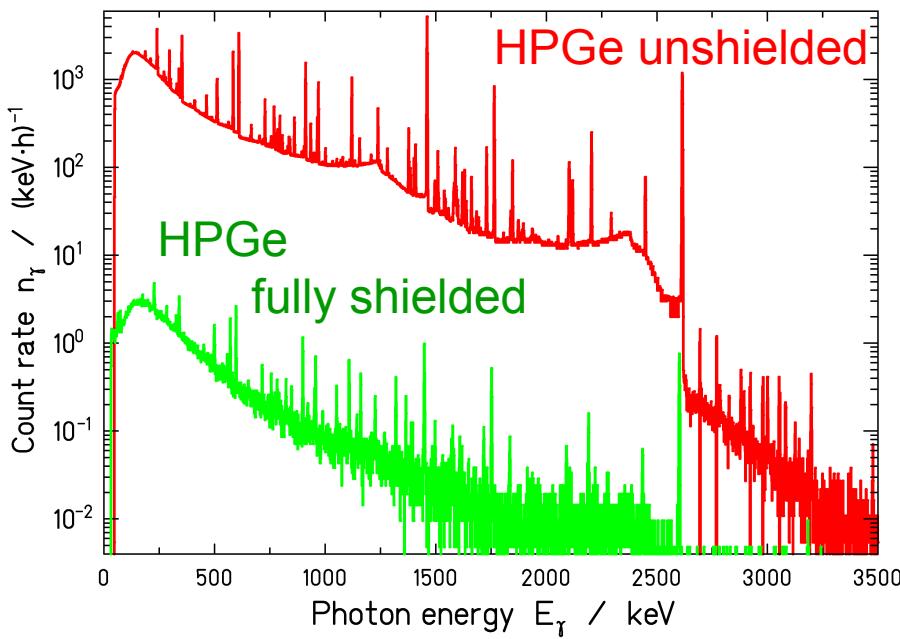


- new target chamber: steel instead of Cu to prevent  $(n,\gamma)$
- recess in target chamber → HPGe det. 4 cm closer to target
- shorter target → better signal-to-noise ratio
- beam tube ( $18 \times 18 \text{ mm}^2$ ) inserted into gas target to limit free path of scattered D



# Environmental background

- Pb-castle: to reduce  $^{40}\text{K}$  + U/Th-decay chain  $\gamma$ -lines (low Q-value)
- $\overline{\text{Rn}}$ -box  $\rightarrow$  stable + low background



- 10 cm HD-PE(5% B) shield  $\rightarrow$  reduce n-background to other experiments at LNGS by a factor 5

# Summary and conclusions

- experiment almost ready to start
- $\Phi_n < 10 \text{ n/s}$ :  $\alpha \rightarrow D_2$  @  $p = 0.5 \text{ mbar}$ ,  $I = 130 \mu\text{A}$ ,  
( $R_{01} = 50 \text{ keV}$ )
- 10 d:  $N_{\text{expect}} \approx 570 \text{ cts. assuming}$   
 $S(E_{\text{cm}})$  Mukhamedzhanov (PRC **52** (1995) 3483)
- signal very low → careful investigation of BIB
- $D(^3\text{He}, p)$ -reaction  
( $\sigma$  well known  $\approx \sigma_{D+\alpha}$ , but no signal)
- study implantation  
(saturation effect → impact on BIB)
- beam heating

# Acknowledgements

- to the INFN for the post-doc fellowship
- to the LUNA-collaboration:
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